



Re-evaluation of the ages of $^{40}\text{Ar}/^{39}\text{Ar}$ sanidine standards and supereruptions in the western U.S. using a Noblesse multi-collector mass spectrometer



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ABSTRACT

The rewards from resolving the tempo of magmatic, paleobiologic, paleoclimatic, and tectonic processes at better than the per mil level of precision have spurred a quest within the geochronology community to improve the precision, accuracy, and calibration of several chronometers, including the $^{40}\text{Ar}/^{39}\text{Ar}$ variant of the K-Ar clock. A new generation of multi-collector mass spectrometers for $^{40}\text{Ar}/^{39}\text{Ar}$ geochronology provides the means to address these issues, as well as to evaluate the ages of $^{40}\text{Ar}/^{39}\text{Ar}$ standards and key marker horizons deposited by large volume, caldera-forming supereruptions in the western U.S. at unprecedented temporal resolution. We highlight the utility of a 5-collector Noblesse mass spectrometer that has a sufficiently high mass resolving power to distinguish $^{36}\text{Ar} + \text{H}^{35}\text{Cl}$ from a $^{36}\text{Ar} + \text{H}^{35}\text{Cl} + ^{12}\text{C}_3$ sum peak at m/e 36, maintains a remarkably low and stable background, and thus yields $^{40}\text{Ar}/^{39}\text{Ar}$ dates of exceptionally high precision. We have calibrated a reference gas and developed a simple analytical routine involving one peak hop, both of which are used to evaluate detector-bias correction factors. Because the $^{40}\text{Ar}/^{39}\text{Ar}$ chronometer is a relative dating technique that relies on calibration against mineral standards of known age, precise and accurate determination of the ages of these standards, and verification of isotopic composition and homogeneity, is essential. We present results from both total fusion and incremental heating experiments on single crystals of the widely used Alder Creek (ACs) and Fish Canyon (FCs) sanidine standards. Several recent, independent, measurements of ACs using multi-collector mass spectrometers, including those presented herein, favor an age of $1.1864 \pm 0.0003/0.0012$ Ma (95% confidence analytical/full external uncertainty), that is $>1\%$ younger than the widely used value obtained by Nomade et al. (2005) using a single collector mass spectrometer. Incremental-heating experiments on single crystals of FCs most commonly do not yield plateau ages, instead many crystals appear to be affected by Ar loss or excess Ar. Our findings, obtained using ion counters, independently confirm those recently obtained using a multiple-Faraday detectors or a combination of Faraday and ion counting detectors. Thus, careful consideration must be given to using FCs in future high-precision $^{40}\text{Ar}/^{39}\text{Ar}$ calibrations of the geologic timescale. Notwithstanding, $^{40}\text{Ar}/^{39}\text{Ar}$ dating of sanidine using the 5-collector Noblesse can now provide leverage critical to interpreting the large range of U-Pb zircon dates commonly found in rhyolitic lavas and tuffs.

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1. Introduction

The $^{40}\text{Ar}/^{39}\text{Ar}$ dating technique, pioneered by Merrhue and Turner (1966) nearly 50 years ago, is a versatile geochronometer used to date some of the youngest volcanism on Earth (e.g., Renne et al., 1997) as well as meteorites that formed during the early history of the solar system (Cassata et al., 2010). In large part as a result of the

EARTHTIME initiative that began in 2003, and the recognition of systematic discrepancies between the K-Ar and U-Pb chronometers (Renne et al., 1998; Schoene et al., 2006; Kuiper et al., 2008), demand for improved precision, accuracy, and mass resolution for $^{40}\text{Ar}/^{39}\text{Ar}$ measurements has increased (Schmitz and Kuiper, 2013). Part of this stimulus reflects the rapidly growing use of the U-Pb geochronometer on Phanerozoic rocks. Precise ($\pm 1\%$) and accurate U-Pb ages are possible due to the development chemical abrasion isotope dilution thermal ionization (CA-IDTIMS) methods aimed at eliminating the effects of Pb loss (Mattinson, 2005; Mundil et al., 2004), creation and distribution of the EARTHTIME tracer (Condon et al., 2015), refinement of the isotopic composition of natural uranium (Condon et al., 2010;

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Hiess et al., 2012), and reduction of analytical and systematic uncertainties (Schmitz and Schoene, 2007; McLean et al., 2011, 2015).

Until recently, the vast majority of instruments developed for $^{40}\text{Ar}/^{39}\text{Ar}$ geochronology have been magnetic sector, single collector mass spectrometers that are operated in a peak-hopping mode. During analysis of the gas, the magnetic field strength is changed to direct the various isotopes of Ar onto a single collector. A new generation of multi-collector mass spectrometers for $^{40}\text{Ar}/^{39}\text{Ar}$ geochronology is now becoming more widely used than the traditional single collector instruments (e.g., Mark et al., 2009; Brumm et al., 2010; Rivera et al., 2011; Phillips and Matchan, 2013; Singer et al., 2014; Mercer et al., 2015; Dimaggio et al., 2015). These multi-collector mass spectrometers offer improved precision, higher mass resolution in some cases, and a dramatic reduction in the sample size required for $^{40}\text{Ar}/^{39}\text{Ar}$ analysis. However, detector calibration has hampered routine use for some detector configurations (e.g., Coble et al., 2011).

We have developed a reference gas and a simple analytical routine involving one peak hop to evaluate detector-bias correction factors for our 5-collector Nu Instruments Noblesse mass spectrometer. We re-evaluate the ages of a Cretaceous and a Quaternary supereruption, and build upon several recent attempts to calibrate the ages of two commonly used $^{40}\text{Ar}/^{39}\text{Ar}$ mineral standards, the Alder Creek rhyolite sanidine (ACs; Turrin et al., 1994; Renne et al., 1998; Nomade et al., 2005) and the Fish Canyon Tuff sanidine (FCs; Renne et al., 1998; Kuiper et al., 2008; Rivera et al., 2011). Because the $^{40}\text{Ar}/^{39}\text{Ar}$ method is a relative dating technique, $^{40}\text{Ar}/^{39}\text{Ar}$ ages are referenced to a mineral standard of known age. Thus, precise and accurate determination of $^{40}\text{Ar}/^{39}\text{Ar}$ standard ages is paramount to facilitate comparisons to ages determined using other chronometers, e.g., U–Pb, and to reduce systematic uncertainties. Both total fusion and incremental heating experiments on single crystals of the ACs and FCs were conducted and analyzed using a 5-collector Noblesse mass spectrometer. Single crystal incremental heating experiments are not reported by any of the recent standard calibrations using multi-collector mass spectrometers (Phillips and Matchan, 2013; Gleadow et al., 2015; Rivera et al., 2011, 2013; Coble et al., 2011). Our findings suggest that ACs remains a viable Quaternary $^{40}\text{Ar}/^{39}\text{Ar}$ standard, whereas isotopic heterogeneity in the FCs makes it a non-ideal standard for high-precision timescale studies or chronometer calibrations.

2. Instrumentation

The Nu instruments Noblesse multi-collector mass spectrometer in the WiscAr laboratory at the University of Wisconsin-Madison (model NG022) is equipped with a 75° magnetic sector, Nier-type ion-source, two quadrupole lens arrays, and a fixed collector block (i.e., no moving parts). Volume of the WiscAr Noblesse is $\sim 2300\text{ cm}^3$. The collector configuration features an axial faraday detector and four discrete dynode ETP® ion-counting electron multipliers, two on the high mass side (IC0 and IC1) and two on the low mass side (IC2 and IC3), which allow for simultaneous measurement of all isotopes of Ar.

The ability of the instrument to resolve interferences is described by the mass resolving power (MRP), which is defined as $m/\Delta m$, where m is the mass of the peak and Δm is the mass difference between 5% and 95% of peak height on the side of the peak (Ireland, 2013). The MRP of the WiscAr Noblesse is ~ 3000 , which is considerably higher than that determined in most Noblesse instruments installed prior to 2015 (more recent instruments exceed this MRP; Saxton, 2015). The MRP of the WiscAr Noblesse allows for resolution of hydrocarbons from Ar at m/e 37 to 40; at m/e 36, $^{36}\text{Ar} + \text{H}^{35}\text{Cl}$ can be substantially resolved from combined $^{36}\text{Ar} + \text{H}^{35}\text{Cl} + ^{12}\text{C}_3$ (Fig. 1). To avoid any influence of $^{12}\text{C}_3$ at m/e 36 as well as other isobaric interferences at the other masses, measurements are taken on the low mass side of the argon peaks (Fig. 1). Peak position, width, and shape are optimized by first adjusting

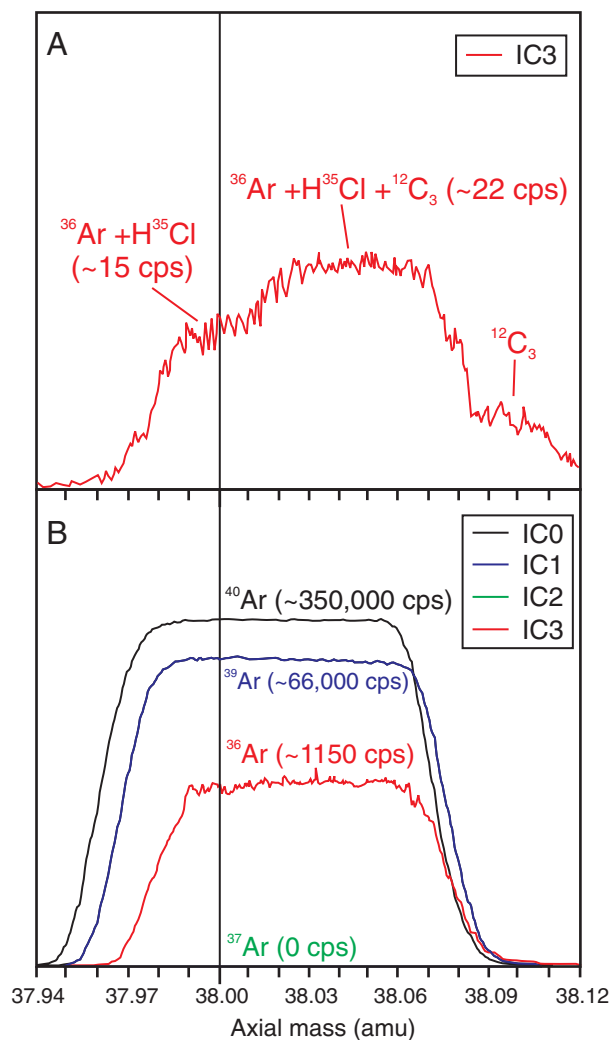


Fig. 1. (A) Typical analytical blank, which includes exposure to getters for 120 s and cryo trap for an additional 60 s. Scan over m/e 36 in IC3 collector showing that $^{36}\text{Ar} + \text{H}^{35}\text{Cl}$ can be partially resolved from a $^{36}\text{Ar} + \text{H}^{35}\text{Cl} + ^{12}\text{C}_3$ sum peak at m/e 36. Y-axis spans 0 to 50 counts per second (cps). (B) A new standard gas developed in February 2015. Scale (in counts per second, cps) for IC0 is 0–400000, IC1 is 0–100000, and IC3 is 0–3000. ^{37}Ar is not present in the standard gas. Solid vertical black line indicates where measurements are made for step 1 of the analytical routine.

the source parameters and then changing the voltages of the quadrupole lenses. The same quadrupole settings are used for measurements of blanks, standard gas, and samples. All analyses in this study were conducted under similar high mass-resolution conditions.

The Noblesse is attached to a gas extraction line with a Photon Machines 60 W CO_2 laser, two SAES GP-50 getters, an ARS cryotrap operating at -125°C , and two gas reservoir/pipette systems. Most analyses conducted are fusions of small mass single sanidine crystals or incremental heating of single sanidine crystals and small aliquots ($<20\text{ mg}$) of volcanic groundmass. For total fusion experiments, the sample is melted in a single step at $\sim 6\text{--}7\text{ W}$ of laser power for 30 s. Gas released from both fusion and incremental heating experiments is cleaned with two SAES GP50 getters, one at 50 W and one at room temperature, for 120 s followed by exposure to the cryotrap for an additional 60 s. Blank and standard gas analyses follow the same gas cleanup routine. Typical analytical blanks are 2500 cps for ^{40}Ar , 40 cps for ^{39}Ar , 20 cps for ^{37}Ar , and 13–15 cps for ^{36}Ar (Fig. 1); these approximately translate to 2.5e^{-17} mol of ^{40}Ar , 4.0e^{-19} mol of ^{39}Ar , 2.0e^{-19} mol of ^{37}Ar , and 1.4e^{-19} mol of ^{36}Ar based on our estimate of instrument sensitivity.

3. Detector linearity

One major limitation of ion-counting electron multipliers is that they are not suited to measure large ion beams. Repeated bombardment with ion beams >500,000 cps degrades IC detectors over time, which requires periodic increase of the voltage applied to the dynode so that it can operate in its plateau region. There is a finite voltage range over which the collector can be operated, and once the maximum collector voltage is reached, it must be replaced. Large ion beams also often result in non-linear behavior of IC detectors (Hoffmann et al., 2005). An IC detector should measure a constant ratio of two isotopes regardless of signal size, provided the detectors respond linearly to input ions and the dead time correction is appropriate. We did not determine the dead time experimentally, but instead employed the isotope ratio approach (e.g., Nelms et al., 2001; Nygren et al., 2011) to evaluate if the instrument dead time setting in the IC0 and IC1 collectors is at an optimum value, and to determine at what signal sizes the instrument dead time setting is no longer appropriate. We measured atmospheric Ar in single collector mode over a range of signal sizes from 50,000 to 875,000 cps (Fig. 2). The intensity range brackets typical ^{40}Ar and ^{39}Ar signals of sample unknowns. For both IC0 and IC1 detectors, the blank corrected $^{40}\text{Ar}/^{36}\text{Ar}$ ratio was plotted against the measured ^{40}Ar intensity (Fig. 2). Both detectors are linear over a very large dynamic range. The slope of the line fit through the data from 50,000 to ~630,000 cps is essentially horizontal (slope $-7\text{e}-7$). When the data from 700,000 to 750,000 cps is included, the slope of the line through the data changes by nearly a factor of 3, and thus the instrument dead time setting is likely no longer appropriate. At 875,000 cps, counts are undoubtedly missed by the IC0 and IC1 detectors, resulting in lower measured $^{40}\text{Ar}/^{36}\text{Ar}$ ratios (Fig. 2). Based on the results of this linearity test, we established a gas splitting routine for signals >630,000 cps, which includes a blank routine that mimics the valve sequence of the gas splitting routine. For the IC2 and IC3

low mass IC detectors, which typically do not exceed 50000 cps of ^{37}Ar and 2000 cps of ^{36}Ar , respectively, the instrument dead time correction for these relatively small signals is likely accurate.

4. Multi-collector calibration

4.1. Standard gas composition

For decades, atmospheric Ar, specifically its $^{40}\text{Ar}/^{36}\text{Ar}$ ratio, has been routinely measured by single collector mass spectrometers outfitted for $^{40}\text{Ar}/^{39}\text{Ar}$ analyses to assess any mass-dependent bias introduced in the ion source or collectors. Coble et al. (2011) outline difficulties in using atmospheric Ar to calibrate source mass discrimination and efficiency contrasts of IC detectors using a 3- collector Noblesse, although these challenges have been surmounted in other laboratories (Rivera et al., 2011; Brumm et al., 2010; Campisano et al., 2014). Our approach to multi-collector calibration is similar to that of Coble et al. (2011) in that we have developed an in-house standard gas to assess mass discrimination and differences in efficiencies of the IC detectors. However, because the collector configuration of the WiscAr Noblesse (4IC + 1 Faraday) is different than that at Stanford (3 IC + 1 Faraday; Coble et al., 2011), we do not need four reference gases. For reasons detailed in Section 4.2, we can simply rely on one standard gas, which contains ^{39}Ar mixed with atmosphere. The ^{39}Ar was produced from fusion of grams of irradiated Fe-doped KAlSiO_4 glass. The ^{39}Ar released from the K-glass was gettered for one hour using the SAES getters on the gas extraction manifold and then exposed to the cryotrap, before expansion into a 5 L reservoir containing atmospheric Ar that is behind a 0.1 cm^3 automated pipette. Our determination of the isotopic composition of the standard gas was done in single collector mode with measurements done on both the IC0 and IC1 detectors. Source mass bias was constrained via measurements of atmospheric Ar in a blank-air-standard routine. The Lee et al. (2006) value for atmospheric Ar was used for standard gas calibration. Standard gas calibration was conducted on both the IC0 and IC1 detectors and measurements were performed periodically over a span of 18 months (Fig. 3). The $^{40}\text{Ar}/^{39}\text{Ar}$, $^{39}\text{Ar}/^{36}\text{Ar}$, and $^{36}\text{Ar}/^{38}\text{Ar}$ are all known to a precision of better than 0.7‰ (Fig. 4). The standard gas shown in Fig. 3 was used for Noblesse measurements made from 2012 through 2014, which includes many of the measurements discussed herein. However, we developed a new standard gas in February 2015 that has a different isotopic composition than its predecessor (Figs. 1B and 3).

4.2. Detector bias correction factors

Analyses of unknowns, blanks, and standard minerals are carried out in identical fashion with a routine involving one peak hop (Table 1). This single routine works for experiments on samples of all ages, which is unlike the complicated, multiple analytical routines outlined by Coble et al. (2011) and Kellett and Joyce (2014) that are dependent on signal size, type of gas analyzed, and sample age. During initial measurement of the gas ^{40}Ar (IC0), ^{39}Ar (IC1), ^{37}Ar (IC2), and ^{36}Ar (IC3) are measured simultaneously, followed by a peak jump of one atomic mass unit where ^{39}Ar (IC0), ^{38}Ar (IC1), and ^{36}Ar (IC2) are measured. This two-step cycle, which takes 35 s, is repeated 15 times. The Faraday detector is not used. Beam switching is achieved by varying the field of the mass spectrometer magnet and with no adjustment to the quadrupole focusing lenses.

The peak hop is necessary to obtain $^{40}\text{Ar}/^{39}\text{Ar}$, $^{36}\text{Ar}/^{39}\text{Ar}$, and $^{37}\text{Ar}/^{39}\text{Ar}$ ratios, which are needed to calculate the age of a sample (McDougall and Harrison, 1999). Accurate and reproducible age determinations using a multi-collector mass spectrometer require that mass fractionation effects and the relative efficiencies of the different detectors be well known. This is achieved by repeating a standard bracketing routine. Analysis of bracketing standard gas aliquots allows

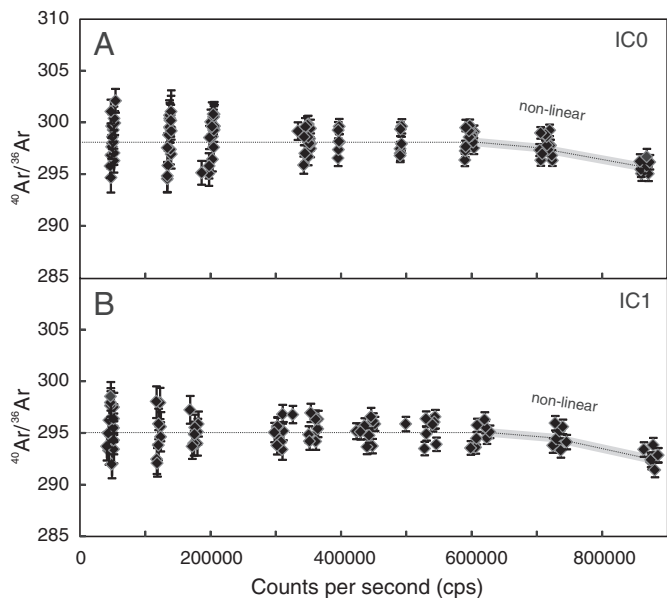


Fig. 2. $^{40}\text{Ar}/^{36}\text{Ar}$ vs ^{40}Ar signal size (cps) for detectors (A) IC0 and (B) IC1. Both IC0 and IC1 detectors exhibit linearity over the range in count rates that are typical of the signals for the samples analyzed in this study. For IC0, the slope of the line is $6\text{e}-7 \times + 298.55$ for signals from 50,000 cps to 400,000 cps and $-7\text{e}-7 \times + 298.77$ for signals from 50,000 cps to 600,000 cps. A change in slope appears to occur with signals greater than ~630,000 cps as the slope of the line fit through 50,000 to 750,000 cps has a slope of $-2\text{e}-6 \times + 298.95$. This suggests that: 1) collector nonlinearity begins to occur with signals >630,000 cps and 2) the instrument dead time correction is no longer appropriate for these large signals.

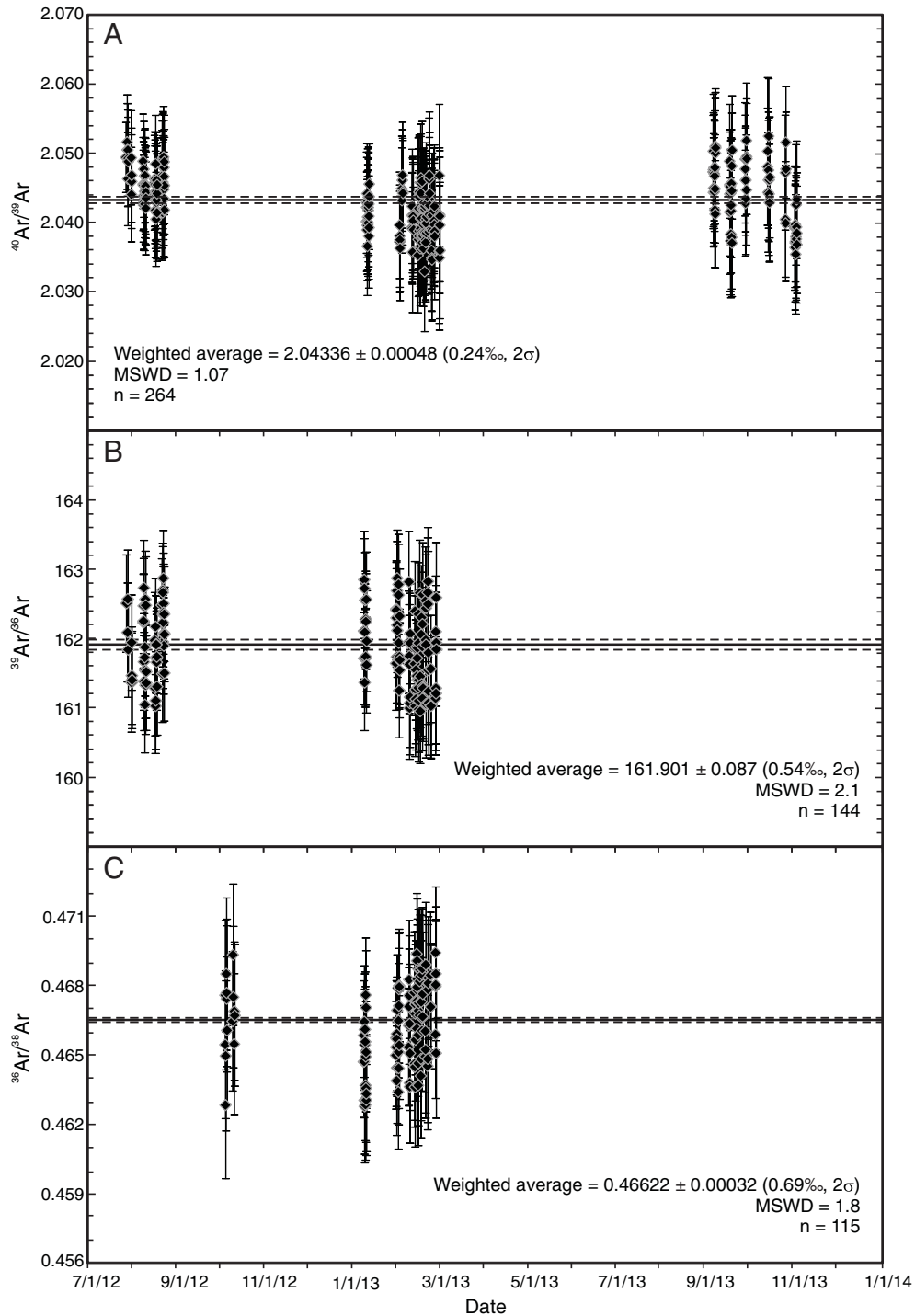


Fig. 3. Standard gas composition as determined by single collector measurements in both IC0 and IC1 collectors. Each standard gas measurement ($\pm 2\sigma$ uncertainty) shown in these panels has been blank and mass bias-corrected. (A) $^{40}\text{Ar}/^{39}\text{Ar}$ (B) $^{39}\text{Ar}/^{36}\text{Ar}$. (C) $^{36}\text{Ar}/^{38}\text{Ar}$. Solid lines represent weighted mean values. Dashed lines are the uncertainties in the weighted mean at the 95% confidence level. We focused on evaluating select isotope ratios ($^{36}\text{Ar}/^{38}\text{Ar}$ in Oct. 2012 and $^{40}\text{Ar}/^{39}\text{Ar}$ in Oct. 2013) at certain points throughout this 18 month period.

the calculation of correction factors, which incorporate mass discrimination of the source and detector, and detector efficiency. The $^{39}\text{Ar}/^{36}\text{Ar}$ is corrected for instrument mass fractionation and detector efficiencies using the following equation:

$$\text{true } ^{39}\text{Ar}/^{36}\text{Ar} (\text{unknown}) = \frac{\text{measured } ^{39}\text{Ar}/^{36}\text{Ar} (\text{unknown})}{[\text{measured } ^{39}\text{Ar}/^{36}\text{Ar} (\text{standard})/\text{true } ^{39}\text{Ar}/^{36}\text{Ar} (\text{standard})]}.$$

The measured $^{39}\text{Ar}/^{36}\text{Ar}$ value for the standard represents the average of the two standard analyses that bracket the sample

unknown. The decay of ^{39}Ar since creation of the standard gas and removal of sample from the reactor are taken into account in all of the calculations.

The measured $^{37}\text{Ar}/^{39}\text{Ar}$ ratio in sample unknowns, where ^{37}Ar is measured in IC2 and ^{39}Ar is measured in IC1, is corrected differently because there is no ^{37}Ar in the standard gas. In this case, ^{36}Ar is measured in IC2 and ^{38}Ar in IC1 in step 2 of the analytical routine so that:

$$\text{true } ^{37}\text{Ar}/^{39}\text{Ar} (\text{unknown}) = \frac{\text{measured } ^{37}\text{Ar}/^{39}\text{Ar} (\text{unknown})}{[\text{measured } ^{36}\text{Ar}/^{38}\text{Ar} (\text{standard})/\text{true } ^{36}\text{Ar}/^{38}\text{Ar} (\text{standard})]}.$$

Table 1

Analysis routine and collector configuration for Noblesse with one Faraday, two high mass, and two low mass ion counting electron multipliers (IC). A typical analytical routine requires one peak hop.

Detector	IC0	IC1	Faraday	IC2	IC3	Integration time
Step 1	^{40}Ar	^{39}Ar	–	^{37}Ar	^{36}Ar	15 s
Step 2	^{39}Ar	^{38}Ar	–	^{36}Ar		20 s

The measured $^{36}\text{Ar}/^{38}\text{Ar}$ value for the standard represents the average of the two standard analyses that bracket the sample unknown. The correction of $^{37}\text{Ar}_{\text{IC2}}/^{39}\text{Ar}_{\text{IC1}}$ using $^{36}\text{Ar}_{\text{IC2}}/^{38}\text{Ar}_{\text{IC1}}$ works even when $^{39}\text{Ar}_{\text{IC1}}$ signals are appreciably larger than the $^{38}\text{Ar}_{\text{IC1}}$ signals. We recognize that this correction assumes that the mass fractionation between ^{37}Ar and ^{39}Ar is the same as that between ^{36}Ar and ^{38}Ar . It should also be noted that a breakdown of the sources of analytical uncertainty assigned to each date in this study indicate that the uncertainty associated with $^{37}\text{Ar}_{\text{IC2}}/^{39}\text{Ar}_{\text{IC1}}$ only contributes on average about 0.6% to the analytical uncertainty versus ~9.3% for $^{40}\text{Ar}_{\text{IC0}}/^{39}\text{Ar}_{\text{IC0}}$ uncertainty and ~90.1% for $^{36}\text{Ar}_{\text{IC3}}/^{39}\text{Ar}_{\text{IC1}}$ uncertainty.

Because both ^{40}Ar and ^{39}Ar are measured in the same detector (IC0), the mass fractionation correction required to convert measured $(^{40}\text{Ar}/^{39}\text{Ar})_{\text{IC0}}$ to the true $(^{40}\text{Ar}/^{39}\text{Ar})_{\text{IC0}}$ is ignored following Brumm et al. (2010), who point out that the effect of this is a small systematic error, but because both monitor and unknown minerals are handled in the same way, the effect cancels out when the ages for the unknowns are calculated. All of the correction factors can vary by up to 2–3 per mil over a 24 h period. However, the variation is typically gradual, and thus is accurately captured by the standard analyses which are done every 45 min. Apart from the dead time correction, which is applied within the Noblesse software, all data reduction (including signal vs. time fits) and corrections (e.g., IC dark noise, ^{37}Ar and ^{39}Ar decay, etc) are carried out offline using an in-house data reduction program.

5. Samples and results

5.1. Alder Creek rhyolite sanidine (ACs)

Turrin et al. (1994) proposed that sanidine from the rhyolite of Alder Creek be used as a Quaternary standard for the $^{40}\text{Ar}/^{39}\text{Ar}$ method, in part because this rock has a transitional direction of magnetization associated with the Cobb Mountain subchron, a period of normal polarity within the Matuyama reversed chron that is recorded globally in marine sediments. Thus, the rhyolite of Alder Creek is an important tie point of for the Geomagnetic Instability Timescale (GITS; Singer, 2014). The most commonly used age for ACs is that of Nomade et al. (2005), which is based on 225 single and multi-crystal total fusion analyses that give a weighted mean of 1.1929 ± 0.0007 Ma (95% confidence interval) relative to a Fish Canyon sanidine standard (FCs) age of 28.02 Ma. The Nomade et al. (2005) age recalculates to 1.2006 ± 0.0007 Ma (Fig. 4), if using a FCs age of 28.201 Ma (Kuiper et al., 2008) (Note: all ACs ages herein will be presented relative to a Fish Canyon sanidine standard age of 28.201 Ma using the Min et al. (2000) decay constant with analytical uncertainties only at the 95% confidence interval unless specified otherwise). Many $^{40}\text{Ar}/^{39}\text{Ar}$ laboratories have attempted to independently determine the age of ACs relative to another standard. The EARTHTIME experiment to intercalibrate laboratories against one another by measuring the isotopic composition of ACs relative to the FCs standard yielded variable results with ACs ages spanning ~3% (gray field in Fig. 4; Heizler and EARTHTIME working group, 2008). Several recent attempts to determine the age of ACs using multi-collector mass spectrometers yield consistent results, despite using different collector calibration routines. Rivera et al. (2013) report that 75 of 88 single crystal fusions of ACs-2, the same sample analyzed by Nomade et al. (2005), give a weighted mean $^{40}\text{Ar}/^{39}\text{Ar}$ age of 1.1864 ± 0.0006 Ma. Coble et al. (2011) present nine dates with an imprecise weighted mean of 1.1877 ± 0.0068 Ma. The youngest 28 $^{40}\text{Ar}/^{39}\text{Ar}$ dates of Phillips and Matchan (2013) give a weighted mean of 1.1847 ± 0.0006 Ma, although this population has a MSWD of 2.7 (Fig. 4). It is important to note that Phillips and Matchan (2013) fused 2–3 crystals per

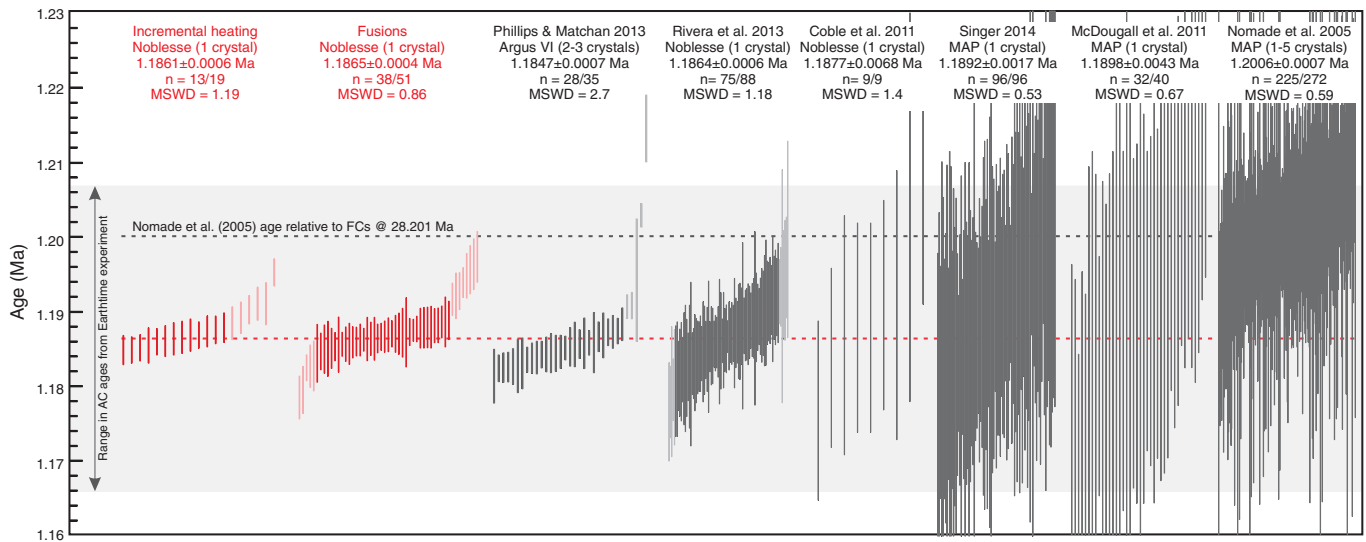


Fig. 4. Comparison of $^{40}\text{Ar}/^{39}\text{Ar}$ ages determined for Alder Creek rhyolite sanidine sample ACs-2 since 2005. $^{40}\text{Ar}/^{39}\text{Ar}$ ages are shown with $\pm 2\sigma$ analytical uncertainties only. Some of the published ages have been recalculated so that all the data in this figure are relative to 28.201 Ma for the Fish Canyon sanidine. The commonly used age from Nomade et al. (2005) is shown as a dotted line. Gray band indicates range in ages generated during the EARTHTIME laboratory intercalibration experiment (Heizler and EARTHTIME working group, 2008). Recent experiments including all the data collected using multi-collector mass spectrometers (Phillips and Matchan, 2013; Rivera et al., 2013) give ages that are significantly younger than the Nomade et al. (2005) age. Our preferred age (red dotted line) for AC-2 based on a weighted mean of the Rivera et al. (2013) dates and our plateau ages and total fusion dates is $1.1864 \pm 0.0003/0.0012$ Ma (95% confidence analytical/full external).

Data sources: Phillips and Matchan (2013); Rivera et al. (2013); Coble et al. (2011); Singer (2014); McDougall et al. (2011); Nomade et al. (2005).

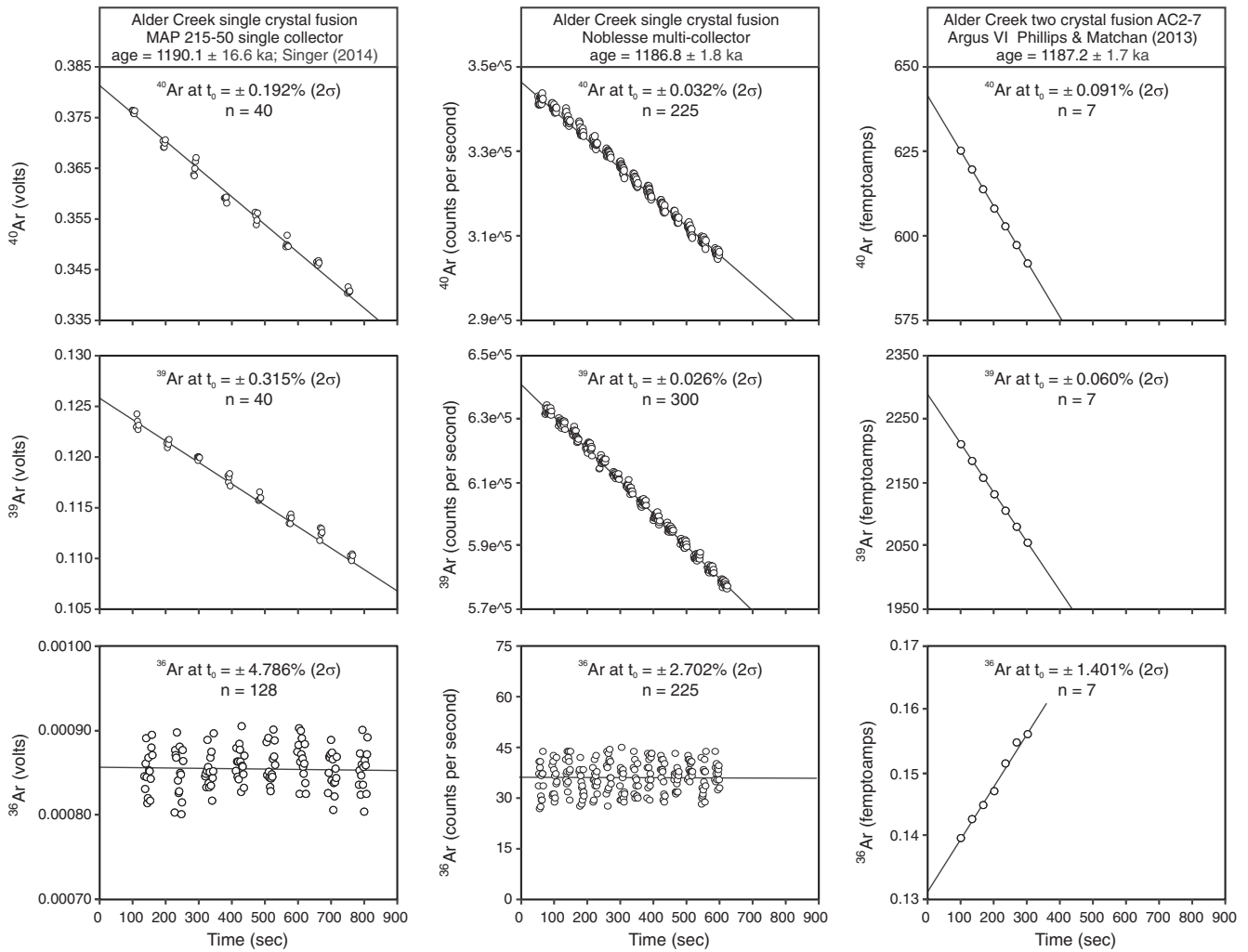


Fig. 5. Comparison of ACS-2 total fusion analysis performed on single collector MAP 215–50 (left column; Singer, 2014), Noblesse (middle column), Argus VI (right column; Phillips and Matchan, 2013). MAP215–50 data collected on a multiplier operated in analogue mode. Integration time for Argus VI data is 33.55 s. Data from instruments is shown before any background or blank correction. Note that whereas there is some reduction in the uncertainty of the ^{36}Ar intercept using the Noblesse, there is marked reduction in the uncertainty associated with the ^{40}Ar and ^{39}Ar intercepts.

analysis, whereas all the other previously mentioned studies performed single crystal fusions.

We employed both single crystal total fusion and single crystal incremental heating experiments to determine the age of the Alder Creek sanidine. All experiments were performed on the ACS-2 standard, which is a split of the same sample analyzed by Nomade et al. (2005) and Rivera et al. (2013). Total fusion dates span 19 ka or ~1.6% from 1.178 ± 0.003 Ma to 1.197 ± 0.003 Ma, which is similar to the range observed by Rivera et al. (2013) and Phillips and Matchan (2013). Thirty-eight of 51 dates give a weighted mean of 1.1865 ± 0.0004 Ma with a MSWD of 0.86 (Fig. 4). Most grains of ACs are highly radiogenic with $^{40}\text{Ar}^*$ typically greater than 90%, but about one-third of the analyzed sanidine had 79–90% $^{40}\text{Ar}^*$. There are tails of $^{40}\text{Ar}/^{39}\text{Ar}$ dates to both the young and old ends of the mean population of 41 dates. A similar trend was observed by Rivera et al. (2013). There is no obvious correlation between age, K/Ca ratio, $^{40}\text{Ar}^*$, or signal size.

The recent development of new generation noble gas mass spectrometers has facilitated considerable focus on improving the precision and accuracy of $^{40}\text{Ar}/^{39}\text{Ar}$ ages, most of which centers on the ^{36}Ar measurements (e.g., Saxton, 2015). Whereas the measurement of ^{36}Ar is critical, especially given the complexities that can now be resolved at m/e 36 (Fig. 1), much of the improved precision obtained using multi-collector mass spectrometers is simply due to improved counting statistics for ^{40}Ar and ^{39}Ar . Fig. 5 compares results from Alder

Creek sanidine analyzed using: (1) the MAP 215–50 and (2) Noblesse mass spectrometers in the WiscAr laboratory, and (3) an Argus VI mass spectrometer (Phillips and Matchan, 2013). Compared to the modest reduction in the uncertainty of the ^{36}Ar intercept when using the Noblesse or Argus VI, uncertainties associated with the ^{40}Ar and ^{39}Ar intercepts are reduced by a factor of ~5–6 using the Noblesse relative to the single-collector MAP 215–50 instrument.

Single crystal incremental heating experiments of ACS-2 were performed as part of an interlaboratory comparison between the WiscAr laboratory, New Mexico Tech, and Oregon State University (Heizler et al., 2015). Nineteen of 26 crystals analyzed using the Noblesse produced plateau ages (73% success rate). Our criteria for an acceptable plateau are similar to those outlined by Sharp and Renne (2005) and Jicha et al. (2012). The youngest 13 plateau ages have a weighted mean of 1.1861 ± 0.0006 Ma, nearly identical to the weighted mean fusion age. Interestingly, none of the 19 plateau ages are younger than 1.1847 ± 0.0017 Ma, but six are slightly older, extending to 1.1951 ± 0.0019 Ma. The origin of the older dates is unclear, but may reflect antecrysts mobilized during eruption of the Alder Creek rhyolite or a uniform distribution of excess Ar during the incremental heating experiment. The integrated ages of the 19 crystals that yielded plateaus span 1.1757 ± 0.0017 Ma to 1.1957 ± 0.0019 Ma, nearly identical to the range exhibited by the total fusions (Fig. 6). The first step of numerous incremental heating experiments is distinctly younger than the plateau

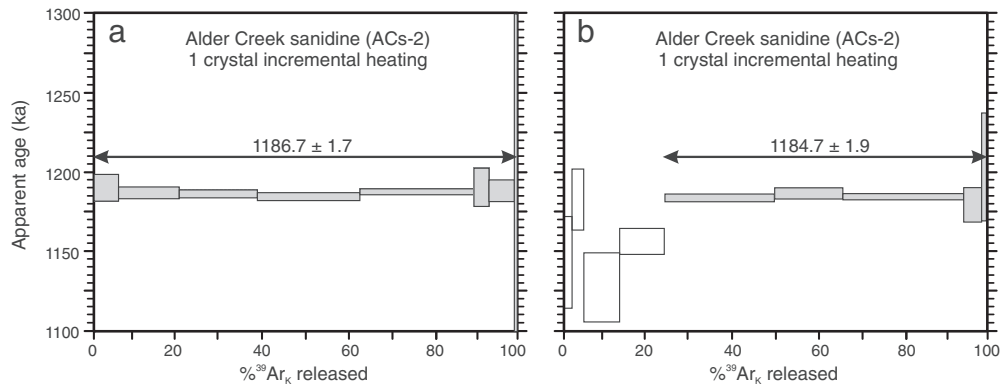


Fig. 6. Representative age spectra from single crystal incremental heating experiments of the Alder Creek sanidine (ACs-2). Ages are calculated relative to the 28.201 Ma Fish Canyon sanidine standard. Most experiments yielded plateaus, some of which a) were 100% concordant, b) give younger ages at the beginning of the experiment (white boxes), or older ages in the higher temperature steps (not shown).

age, which could explain why several total fusion and integrated ages are younger than the youngest plateau ages.

5.2. Fish Canyon Tuff sanidine (FCs)

Sanidine from the Fish Canyon tuff was first proposed as a $^{40}\text{Ar}/^{39}\text{Ar}$ standard by Cebula et al. (1986). It is the most widely used $^{40}\text{Ar}/^{39}\text{Ar}$ mineral standard (Renne et al., 1998), yet its age remains the subject of scrutiny. The FCs age currently used by $^{40}\text{Ar}/^{39}\text{Ar}$ laboratories ranges from 27.64 to 28.29 Ma (Hildreth et al., 2014; Renne et al., 2011). $^{238}\text{U}/^{206}\text{Pb}$ zircon dates from the Fish Canyon Tuff vary from 28.0 to 28.6 Ma (e.g. Schmitz and Bowring, 2001; Bachmann et al., 2007; Rivera et al., 2011; Wotzlaw et al., 2013). Notably, Wotzlaw et al. (2013) used the CA-ID-TIMS method to determine $^{238}\text{U}/^{206}\text{Pb}$ dates from 24 individual zircons in the Fish Canyon tuff, finding that zircon records crystallization of the magma body for 442 kyr ending with the youngest zircon measured at 28.196 ± 0.038 Ma (2σ analytical). Several studies over the last five years have attempted to reduce the systematic uncertainties associated with the $^{40}\text{Ar}/^{39}\text{Ar}$ method by calibrating

the age of FCs. Kuiper et al. (2008) cross-calibrated $^{40}\text{Ar}/^{39}\text{Ar}$ sanidine ages from tephras in the astronomically tuned Messâdit section (Melilla Basin, Morocco) to determine the age of FCs as $28.201 \pm 0.012/0.046$ Ma (2σ , analytical/external). Rivera et al. (2011) used a similar technique to cross calibrate A1 tephras sanidines (A1Ts) of the astronomically tuned Faneromeni section (Crete), resulting in an FC age of $28.172 \pm 0.018/0.028$ Ma (2σ , analytical/external). Renne et al. (2010, 2011) used a different approach to obtain a FCs age of 28.294 ± 0.072 Ma (2σ , external) by jointly determining the $^{40}\text{Ar}^*/^{40}\text{K}$ for FCs and ^{40}K decay constants based on published ^{40}K activity and $^{40}\text{Ar}/^{39}\text{Ar} + ^{238}\text{U}/^{206}\text{Pb}$ data pairs (Renne et al., 2011). The Kuiper et al. (2008) FCs age is indistinguishable at 2σ from the astronomically calibrated FCs age of Rivera et al. (2013), the FCs age of Renne et al. (2011), and the youngest zircon in the Fish Canyon Tuff dated by Wotzlaw et al. (2013). Two recent studies used the Argus VI multi-collector mass spectrometer with Faraday detectors at Melbourne University to produce FCs ages that are significantly more precise than each of the previously published FCs ages. Phillips and Matchan (2013) performed 30 laser fusions on 2–4 crystal aliquots and obtained

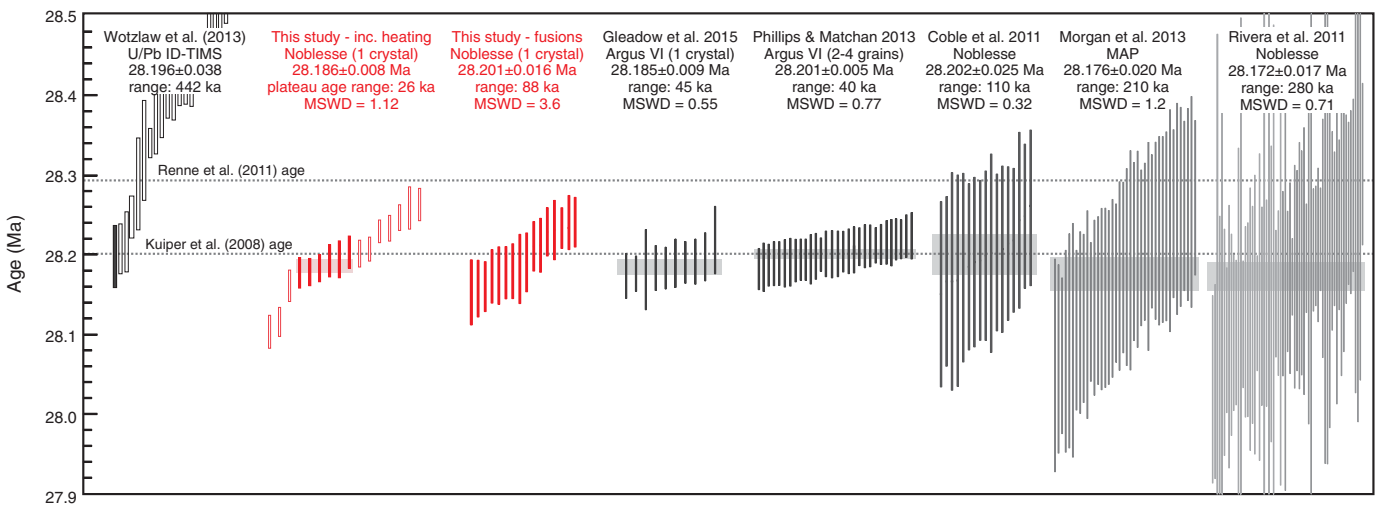


Fig. 7. Comparison of $^{40}\text{Ar}/^{39}\text{Ar}$ ages determined for Fish Canyon sanidine sample FC-2. All $^{40}\text{Ar}/^{39}\text{Ar}$ ages are shown with analytical uncertainties only and are relative to 28.201 Ma for the Fish Canyon sanidine (Kuiper et al., 2008). Note that all total fusion data were performed on single crystals except for that of Phillips and Matchan (2013), which fused 2–4 crystals per analysis. Our single crystal total fusion experiments gave a range in dates spanning ~88 ka and thus have a MSWD of 3.6. Overall the range in dates is similar to those of the previous studies. Only 6 of the 16 single crystal incremental heating experiments gave statistically acceptable plateaus (filled red boxes). A weighted mean age of these 6 plateaus is 28.186 ± 0.008 Ma. The total fusion ages for the crystals which did not yield plateaus are shown as open boxes. CA-IDTIMS U–Pb zircon dates are from Wotzlaw et al. (2013); the youngest zircon (solid black bar) is 28.196 ± 0.038 Ma. We recognize that there are other published U–Pb zircon datasets for the Fish Canyon Tuff. However, the purpose of the figure is to highlight the age complexity within the FCT not select an absolute age for the FCT.

Data sources: Wotzlaw et al. (2013); Gleadow et al. (2015); Phillips and Matchan (2013); Coble et al. (2013); Morgan et al. (2013); Rivera et al. (2011).

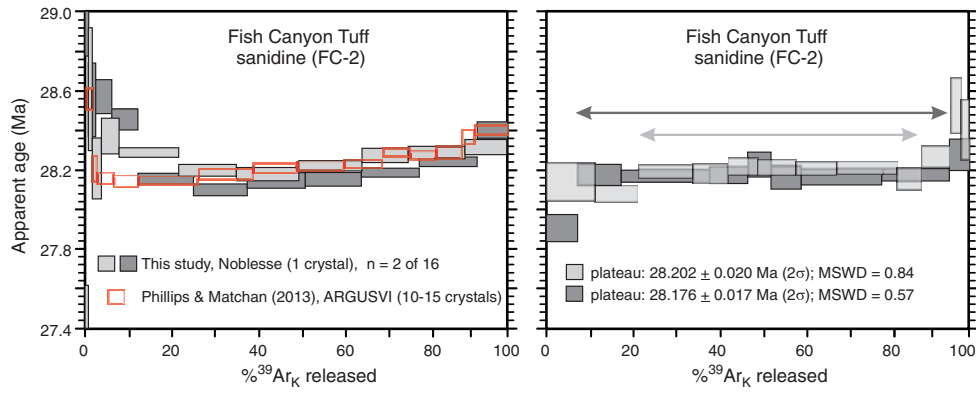


Fig. 8. Age spectrum diagrams for Fish Canyon sanidine (FC-2). Ages are calculated relative to an age of 28.201 Ma for FCs. Several of the discordant age spectra show a gas release pattern that is similar to those observed by Phillips and Matchan (2013) (left). Only 6 of the 16 single crystal incremental heating experiments gave statistically acceptable plateaus, two of which are shown (right).

an age of 28.201 ± 0.005 Ma, whereas 10 single crystal fusions of Gleadow et al. (2015) give a slightly younger FCs age of 28.185 ± 0.009 Ma. Both of the Melbourne University FCs datasets used Fish Canyon sanidine as the fluence monitor, and therefore cannot truly be used to assess the age of FCs.

Sixteen single crystal ($< 100 \mu\text{m}$) fusions analyzed using the WiscAr Noblesse multi-collector mass-spectrometer yield $^{40}\text{Ar}/^{39}\text{Ar}$ dates that span 88 ka and have a MSWD of 3.6. This aliquot of FCs was monitored with the 1.186 Ma Alder Creek sanidine standard (Rivera et al., 2013) and is from a different irradiation than the one used to determine the age of the Alder Creek sanidine standard (Fig. 4). All FCs used in this study come from the split FC-2, which was collected in 1994, separated, and distributed by New Mexico Tech. The average uncertainty on the individual FCs dates using the Noblesse is ± 30 ka. FCs dates with smaller uncertainties of ± 10 ka have been obtained using an Argus VI mass spectrometer because the ^{40}Ar signal can be directed into a Faraday detector (Heizler, 2012), whereas the ^{40}Ar signals for the IC0 collector of the Noblesse are limited to $\sim 650,000$ cps (Fig. 3). In an attempt to understand the age dispersion among the fusion results, we incrementally heated 16 single FCs crystals from two different irradiations. One aliquot of FCs was monitored with the 28.52 Ma Taylor Creek sanidine standard and the other was monitored with FCs itself. We recognize that by monitoring FCs with FCs and Taylor Creek sanidine, whose age has been determined relative to Fish Canyon sanidine (Renne et al., 1998), we cannot weigh in the ongoing debate as to which FCs age calibration is the most accurate. In order to independently determine the age of FCs, one needs to measure FCs as an unknown against a standard whose age has been independently established by other means as in the case of the astronomically dated 6.943 ± 0.005 Ma (2σ) A1 tephra (A1Ts; Kuiper et al., 2004; Rivera et al., 2011) an astronomically-calibrated $^{40}\text{Ar}/^{39}\text{Ar}$ standard such as the and Alder Creek sanidine (Rivera et al., 2013), or a primary standard such as GA-1550 biotite (Renne et al., 1998).

Only six of the 16 single crystal incremental heating experiments gave a statistically acceptable age plateau (37.5% success rate compared to 73% for the ACs) (Fig. 7). A weighted mean age of these 6 plateaus is 28.186 ± 0.008 Ma, which is nearly identical to the weighted mean age of Gleadow et al. (2015) of 28.185 ± 0.009 Ma. Several of our single crystal incremental heating experiments produced saddle shaped or rising gas release patterns, similar to the multi-crystal (10–25 crystals) experiments reported in Phillips and Matchan (2013) and Gleadow et al. (2015) (Fig. 8). The integrated ages for the sanidine grains that did not form a plateau are both younger and older than the six plateau ages and span a larger range than that observed during the total fusions. Most of the integrated ages are older than 28.2 Ma (Fig. 7), and may reflect the incorporation of excess Ar into the crystal structure.

5.3. Lava Creek tuff sanidine

In addition to investigating the ages and argon systematics of commonly used $^{40}\text{Ar}/^{39}\text{Ar}$ standards, we evaluate the ages of key marker horizons from supereruptions in the western U.S. The Lava Creek Tuff ($> 1000 \text{ km}^3$) is the youngest of the three Yellowstone caldera-forming supereruptions. Christiansen (2001) identified two Lava Creek Tuff members, A and B, which are separated from each other by a bed of fallout. Proximal member B ignimbrite is comparable in volume to member A, but the distal B tephra blanketed much of the central and western United States. Its occurrence in numerous sedimentary and pluvial lake bed sequences makes it a key Quaternary chronostratigraphic marker.

The Lava Creek Tuff member B sample was collected by Tiffany Rivera from a road cut along state highway 47, south of the Lower Mesa Falls turn off ($44^\circ 08.446$, $-111^\circ 17.405$, 1760 m). It corresponds to the sample site of Bindeman and Valley (2001) and is from the more densely welded part of the unit. The sample is light gray to purple with large phenocrysts of sanidine (3–5 mm) and quartz (1–3 mm).

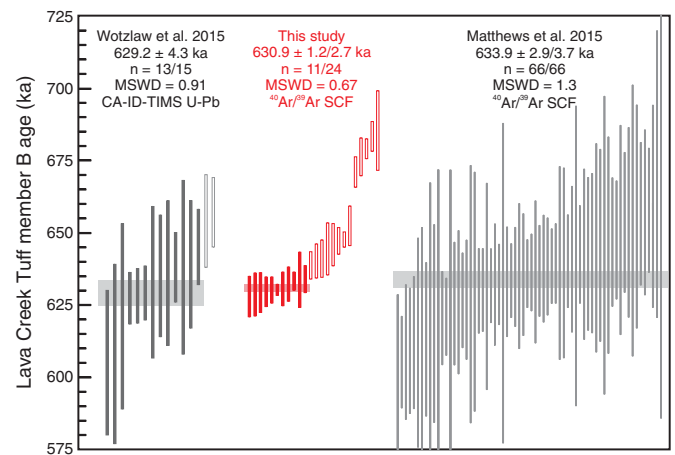


Fig. 9. Comparison of $^{40}\text{Ar}/^{39}\text{Ar}$ sanidine total fusion dates from Lava Creek Tuff member B from Matthews et al. (2015) and this study to the CA-ID-TIMS U–Pb dates of Wotzlaw et al. (2015). The small analytical uncertainties generated using the Noblesse allow us to distinguish between antecrysts that are a little as 8 kyr older than the juvenile population of dates that give a weighted mean age of $630.9 \pm 1.2/2.7$ ka. Uncertainties are given at 95% confidence level, and reported in the form $\pm X/Y$ where X is the analytical uncertainty and Y includes systematic or full external uncertainties. External error propagation methods follow the algorithms of Kuiper et al. (2008). Horizontal boxes represent range of analytical uncertainty for each dataset. SCF = single crystal fusion. For complete data, see Rivera and Jicha, 2015 or electronic supplement.

Pumice distribution within the unit is somewhat bimodal, with typical pumice lengths being ~1 cm or 5–10 cm, all of which are flattened such that the length to width ratio is 10:1.

Previous K–Ar and $^{40}\text{Ar}/^{39}\text{Ar}$ dating of sanidine (Obradovich, 1992; Gansecki et al., 1998; Lanphere et al., 2002) from the Lava Creek Tuff members A and B yield dates ranging from 560 to 700 ka. A recent study by Matthews et al. (2015) found that Lava Creek Tuff member B sanidine yielded a unimodal population of 66 single crystal $^{40}\text{Ar}/^{39}\text{Ar}$ dates with a weighted mean of $634.0 \pm 2.9/3.7$ ka, and an isochron date of $631.9 \pm 3.5/4.3$ ka. The $^{40}\text{Ar}/^{36}\text{Ar}$ intercept of the isochron is higher than, albeit barely overlapping with the atmospheric value when considering the associated uncertainties, thereby suggesting the presence of excess argon. Thirteen of 15 CA-ID-TIMS U–Pb zircon dates from Lava Creek B give a weighted mean of 629.2 ± 4.3 ka (Wotzlaw et al., 2015).

Analyses of 24 single crystal sanidine fusions from Lava Creek Tuff member B using the Noblesse give dates ranging from 627.8 ± 7.0 ka to 685.2 ± 13.7 ka (Fig. 9; Rivera and Jicha, 2015), which is consistent with previous single crystal fusion data (Gansecki et al., 1998; Matthews et al., 2015). The youngest 11 dates have a weighted mean age of $630.9 \pm 1.2/2.7$ ka, which we interpret as our best estimate of the eruptive age. The $630.9 \pm 1.2/2.7$ ka age is within error of the $634.0 \pm 2.9/3.7$ ka $^{40}\text{Ar}/^{39}\text{Ar}$ age of Matthews et al. (2015) and the U–Pb age of 629.2 ± 4.3 ka age of Wotzlaw et al. (2015). It is important to note that the Matthews et al. (2015) age is slightly older because their weighted mean age includes dates that they suggest come from crystals that contain excess argon. However, because the dates generated by the Noblesse have significantly smaller analytical uncertainties (average of ± 6 ka) than those of the Matthews et al. (2015) (average of ± 28 ka), we are able to exclude dates that are as little as 8 ka older (i.e., 638.6 ± 4.7 ka) than the youngest mean group of dates (630.9 ± 1.2 ka). Single crystal incremental heating is required in order to determine whether the older dates reflect crystals that contain excess argon or are antecrysts such as those found in the Huckleberry Ridge Tuff (Singer et al., 2014).

5.4. Cretaceous B bentonite

The distal ash falls from four voluminous eruptions near the Cenomanian–Turonian boundary (ca. 94 Ma) are preserved as bentonites that can be traced through much of the western interior of the United States (Elder, 1988). Bentonite marker bed B is the largest and most widely traceable of these four marker horizons, with deposits found in at least 11 U.S. states and in Alberta, Canada (Elder, 1988). It was deposited during the heavy carbon isotopic excursion termed Ocean Anoxic Event 2 (Schlanger and Jenkyns, 1976) within the *Neocardioceras juddii* ammonite biozone. The isopach map of bed B suggests a source area near the U.S.–Canada border in the region of northeastern Washington/northwestern Montana (Elder, 1988). Although the B bentonite is ~90 cm near the Montana border, it is 135 cm thick in the Alberta foothills, which suggests origin from a source further north (Barker et al., 2011).

We analyzed sanidine from samples K-07-01B (Meyers et al., 2012) and 90-O-19 (Meyers et al., 2012; Obradovich, 1993). K-07-01B corresponds to Bed 80 from GSSP type locality in Pueblo, CO (Kennedy et al., 2005) and was collected by B. Sageman & D. Condon. Sample 90-O-19 is a 15–20 cm white bentonite from Thayer County, Nebraska that corresponds to sample location 20 in Obradovich (1993). Eighteen single sanidine fusions of K07-01B give dates ranging from 93.74 ± 0.11 Ma to 94.72 ± 0.08 Ma (Fig. 10; Supplementary material). The oldest three and youngest two dates are excluded from this group because they have a normalized median absolute deviation (nMAD) greater than 1.5 and are considered outliers (Powell et al., 2002). The weighted mean age of the remaining 13 dates is 93.99 ± 0.03 Ma (Fig. 10). For 90-O-19, 13 of 16 dates give a weighted mean age of 93.98 ± 0.03 Ma, which is nearly identical to our new $^{40}\text{Ar}/^{39}\text{Ar}$ age of K07-01B and within error of the single and multi-crystal fusion $^{40}\text{Ar}/^{39}\text{Ar}$ ages of 94.10 ± 0.37 Ma and 94.07 ± 0.31 Ma from Meyers et al. (2012) for K07-01B and 90-O-19, respectively. Meyers et al. (2012) suggested that the U–Pb age of $94.01 \pm 0.04/0.14$ Ma from the B bentonite in Thayer County, Nebraska (sample site NE-08-01 = 90-

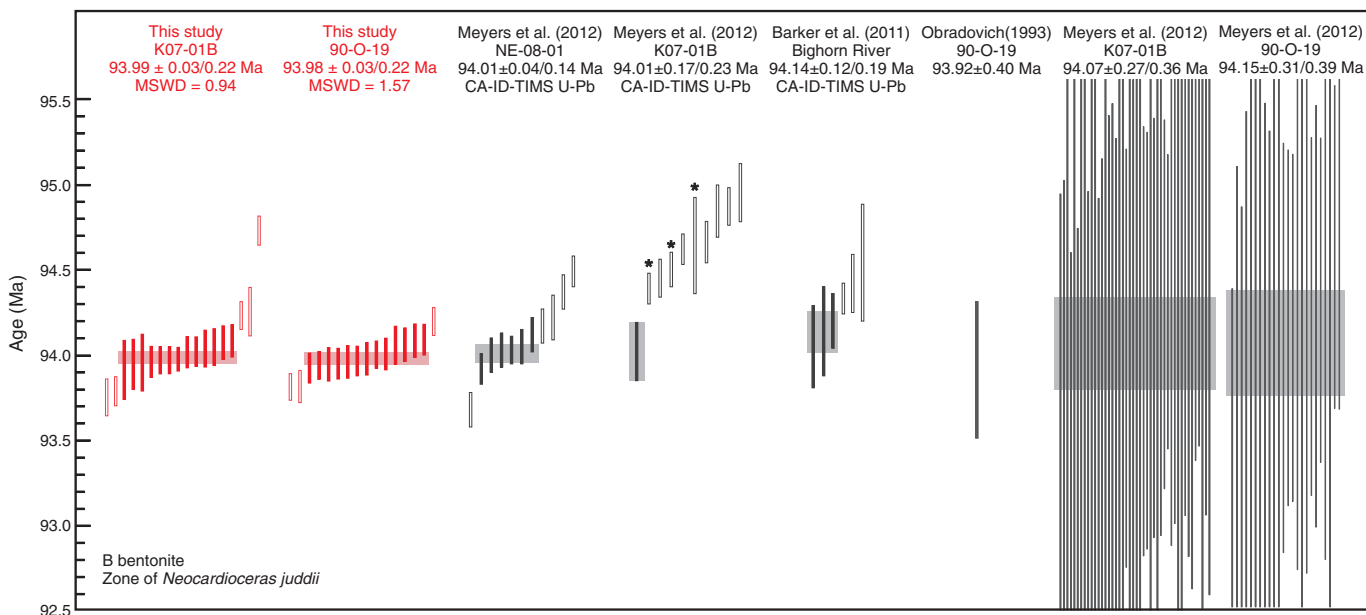


Fig. 10. Compilation of all $^{40}\text{Ar}/^{39}\text{Ar}$ and U–Pb data for the B bentonite of Elder (1988). The B bentonite was sampled at the GSSP Pueblo, Colorado (K07-01B) and in Thayer County, Nebraska (sample 90-O-19 of Obradovich (1993); and NE-08-01 of Meyers et al. (2012) come from the exact same outcrop). The Bighorn River bentonite in Alberta, Canada is interpreted to be the equivalent of the B bentonite (Barker et al. (2011) and references therein). $^{40}\text{Ar}/^{39}\text{Ar}$ multi-crystal sanidine total fusion dates from Meyers et al. (2012) were obtained on a MAP 215–50 single collector instrument versus single crystal total fusions using the Noblesse multi-collector mass spectrometer in this study. Crystals from the same sanidine separate were analyzed in both studies. Youngest U–Pb zircon date from K07-01B is in agreement with the $^{40}\text{Ar}/^{39}\text{Ar}$ ages. Asterisk indicates the U–Pb dates used in Meyers et al. (2012) for this sample, which gave an age of 94.44 ± 0.15 Ma. Youngest three U–Pb dates from Barker et al. (2011) give an age of 94.14 ± 0.12 Ma (MSWD = 0.55), which is within uncertainty of both $^{40}\text{Ar}/^{39}\text{Ar}$ and U–Pb ages for the B bentonite in this study and in Meyers et al. (2012). The U–Pb age of 94.29 ± 0.13 Ma (MSWD = 2.4) as reported by Barker et al. (2011) based on all six zircon dates is older than the all $^{40}\text{Ar}/^{39}\text{Ar}$ and U–Pb ages for the B bentonite. Uncertainties are reported in same manner as Fig. 9.

0–19) is the most accurate estimate of time since eruption, whereas the K-07–01B sample contains a large proportion of inherited zircons. The U–Pb age of $94.01 \pm 0.04/0.14$ Ma is in good agreement with the new high precision $^{40}\text{Ar}/^{39}\text{Ar}$ age of $93.99 \pm 0.03/0.22$ Ma and $93.98 \pm 0.03/0.22$ Ma for the B bentonite. Interestingly, the youngest zircon date of $94.01 \pm 0.17/0.23$ Ma from K07–01B is also in agreement with the new $^{40}\text{Ar}/^{39}\text{Ar}$ dates and the U–Pb age from Nebraska. The new, high-precision age for the B bentonite based on the $^{40}\text{Ar}/^{39}\text{Ar}$ and the U–Pb ages from Colorado and Nebraska is $93.994 \pm 0.018/0.093$ Ma. The Bighorn River bentonite in Alberta, Canada is interpreted to correlate with the B bentonite (Barker et al., 2011 and references therein). A U–Pb age of $94.29 \pm 0.13/0.20$ Ma (MSWD = 2.4) (Barker et al., 2011), based on six zircon dates, is slightly older than our new, high-precision age if only analytical uncertainties are considered. However, the youngest three zircons give an age of $94.14 \pm 0.12/0.19$ Ma (MSWD = 0.55) that is within error of the high-precision $^{40}\text{Ar}/^{39}\text{Ar}$ age. $^{40}\text{Ar}/^{39}\text{Ar}$ ages with the precisions obtained using the Noblesse now enable improved interpretation of the ranges of dates produced using CA-IDTIMS U–Pb zircon dating.

6. Discussion

6.1. Age of the Alder Creek sanidine standard

The Alder Creek sanidine has somewhat less age complexity when compared to the Fish Canyon sanidine age distributions (Figs. 4, 6, 7). Incremental heating (IH) experiments of a few ACs crystals gave saddle shaped spectra like those observed by Phillips and Matchan (2013), but the majority of the IH experiments on single ACs crystals (73%) yielded plateaus (Heizler et al., 2015). Despite the fact that the radiogenic ^{40}Ar yield can be variable and some of the crystals appear to contain excess Ar, the ACs appears to be a viable Quaternary $^{40}\text{Ar}/^{39}\text{Ar}$ standard. Rivera et al. (2013) recommended use of a $^{40}\text{Ar}/^{39}\text{Ar}$ age of $1.1864 \pm 0.0006/0.0016$ Ma (95% confidence analytical/full external) for ACs, but this has not been widely adopted. All of the recent single crystal age determinations of sanidine from the rhyolite of Alder Creek (ACs) that employ multi-collector mass spectrometry including the findings of Rivera et al. (2013) as well as both single crystal total fusion and incremental heating results presented here, results in an age of $1.1864 \pm 0.0003/0.0012$ Ma (95% confidence analytical/full external). This age should now serve as the preferred age for ACs because it is more accurate than the commonly used age of Nomade et al. (2005), which is more than 1% older. The origin of age discrepancy is unknown, but it is likely related to analytical conditions during the Nomade et al. (2005) study. These may include inaccurate measurement of the ^{36}Ar blank and its $^{40}\text{Ar}/^{36}\text{Ar}$ ratio, or improper assignment of time zero, the time at which the gas is expanded from the extraction system into the mass spectrometer.

The transitionally magnetized rhyolite of Alder Creek is the type locality for the Cobb Mountain normal polarity subchron. Rivera et al. (2013) suggest that the astronomical ages of 1.173 and 1.185 Ma for the Cobb Mountain subchron determined using the sediment record of Horng et al. (2002) are the most accurate, and thus the Alder Creek rhyolite records the older reversal at the base of the subchron. However, to our knowledge, there are no other known sediments or lavas globally, which potentially record a 1.173 Ma reversal bounding the Cobb Mountain subchron. In cores from ocean drilling program sites 983 and 984 in the North Atlantic Ocean characterized by high deposition rates, Channell et al. (2002) found the Cobb Mountain subchron to be defined by abrupt polarity transitions at its base and top associated with astronomical ages of 1.215 and 1.190 Ma, respectively. Moreover, the PISO-1500 stack of Channell et al. (2009) that is based on 14 high resolution relative paleointensity and $\delta^{18}\text{O}$ records reveals that the Cobb Mountain subchron is associated with a 30 kyr period of exceptionally low intensity between 1215 and 1185 ka. Thus, a $1.1864 \pm 0.0003/0.0012$ Ma age for ACs argues that the transitionally magnetized

rhyolite of Alder Creek records a snapshot of the younger reversal at the top of the Cobb Mountain subchron, not the older reversal at its base (Singer, 2014). This is further supported by a 1.211 ± 0.011 Ma age determined for a transitionally-magnetized lava flow at Persani Mountain, Romania that is interpreted to record the reversal at the base of this subchron (Panaiotu et al., 2013).

It is important to note that if a 1.1864 Ma age for ACs is adopted, several $^{40}\text{Ar}/^{39}\text{Ar}$ age discrepancies for prominent Quaternary eruptions are eliminated. For example, the age of the Bishop Tuff remains controversial as there are significant differences between published $^{40}\text{Ar}/^{39}\text{Ar}$ dates, ranging from 767.3 ± 3.0 ka (Zeeden et al., 2014) to 780 ± 2 ka (Simon et al., 2014). Numerous recent age constraints based on other chronometers including astrochronology (Zeeden et al., 2014), U–Pb ID-TIMS (Crowley et al., 2007; Ickert et al., 2015), and U–Pb SIMS (Chamberlain et al., 2014a; Ickert et al., 2015) favor a Bishop Tuff eruption age of ~763–773 ka, however, there are some assumptions associated with each of these age determinations (see Ickert et al., 2015 for discussion). If the $^{40}\text{Ar}/^{39}\text{Ar}$ data of Simon et al. (2014) are recalculated using a 1.1864 Ma age for ACs and the Min et al. (2000) decay constant, their age for the Bishop Tuff becomes 768 ± 2 ka, which is in agreement with all other recent age determinations.

6.2. Age heterogeneity in the Fish Canyon sanidine

For reasons mentioned in Section 5.2, the primary goal of our single crystal fusion and incremental heating experiments of the FCs was not to arrive at an independent age for FCs, but rather to evaluate its viability as a $^{40}\text{Ar}/^{39}\text{Ar}$ standard. However, our FCs used for the total fusion experiments was monitored with the astronomically calibrated Alder Creek sanidine (Rivera et al., 2013). The youngest 10 fusions give a weighted mean age of 28.175 ± 0.0011 Ma, which, at a first-order, (1) verifies the $28.172 \pm 0.017/0.028$ Ma astronomical calibration of Rivera et al. (2011), (2) suggests that the $R_{\text{FC}/\text{AC}}$ of the Quadlab and WiscAr laboratory is nearly identical, and (3) indicates that the 28.294 Ma age recommended by Renne et al. (2011) is too old. A much larger dataset set relative to an astronomically calibrated standard is required to more quantitatively evaluate the age of FCs.

A comparison of the FCs spectra from our single crystal incremental heating experiments to those produced in several published multi-crystal step heating experiments (e.g., Spell and McDougall, 2003; Bachmann et al., 2007; Phillips and Matchan, 2013; Gleadow et al., 2015) indicates that the gas release patterns can be reproduced at the single grain level (Fig. 7). It should not come as a surprise that age complexity would be present in the in FCs because it comes from a voluminous ash-flow tuff associated with a supereruption. It is well known that silicic tuffs may contain alkali feldspars much older than their depositional age, and that these reflect inheritance from older rocks at or below the earth's surface (e.g., Lo Bello et al., 1987; Deino and Potts, 1990). Remarkably, advances in both U–Pb and $^{40}\text{Ar}/^{39}\text{Ar}$ methods have begun to reveal that many supereruption deposits contain not only ancient inherited xenocrysts, but also antecrysts and remobilized, partly solidified, portions of magma chambers that include both sanidine and zircon slightly older than the age of eruption (e.g., Chamberlain et al., 2014a, 2014b; Rivera et al., 2014; Wotzlaw et al., 2013). Renne et al. (2012) demonstrated that inherited ^{40}Ar , such as that in antecrysts or remobilized roof rocks, can be retained in alkali feldspars despite immersion in magma at ~900 °C for tens to >100 years, thereby offering a mechanism to explain some of the slightly older ages in the FCs.

We concur with the conclusion of Phillips and Matchan (2013) that FCs is a non-ideal standard for high precision $^{40}\text{Ar}/^{39}\text{Ar}$ geochronology. Thus, an urgent short-term goal for the $^{40}\text{Ar}/^{39}\text{Ar}$ community must be to identify and calibrate a more reliable sanidine standard, preferably one from a distal airfall deposit or small-volume rhyolitic lava or dome of similar age to FCs. Accurate and precise determination of a new standard is needed because analytical uncertainties from multi-

collector mass spectrometers are now significantly smaller than the full external uncertainties.

6.3. Implications for the Cretaceous timescale

As demonstrated by our sanidine total fusion data from the B bentonite of Elder (1988), we can now obtain $^{40}\text{Ar}/^{39}\text{Ar}$ dates from individual Cretaceous sanidine crystals with analytical uncertainties of <0.03 Ma, which rival the uncertainties of the most precise ID-TIMS U–Pb zircon dates for this time period. One drawback of the new, high-precision $^{40}\text{Ar}/^{39}\text{Ar}$ data is that age distributions are more complex because dates often span a continuum (Fig. 10), like many of the ID-TIMS U–Pb zircon datasets. The practice of calculating a weighted mean eruptive age from a subgroup of the youngest crystals can be compromised due to the fact that some crystals may exhibit Ar loss. In some cases, there may not be a single “age” that can be gleaned from a distribution. What is clear is that rigorous statistical interrogation/filtering of high-precision $^{40}\text{Ar}/^{39}\text{Ar}$ and U–Pb data are now required to identify an eruptive/depositional age. The greatly improved precision and accuracy of Cretaceous $^{40}\text{Ar}/^{39}\text{Ar}$ and U–Pb ages now allow us to: (1) confirm or consider alternative correlations of some ash beds in the Western Interior Basin, specifically those of supereruptions with associated tephra fall deposits that span vast regions, (2) re-examine stratigraphic data to potentially identify previously obscured hiatuses, (3) precisely constrain sedimentation rates, and (4) anchor floating astronomical time scales for portions of the Cretaceous that have a resolving power better than the duration of a 100 kyr eccentricity cycle and can thus be used to pinpoint rapid environmental fluctuations and biological changes.

7. Conclusions

- The high mass resolving power of the Noblesse allows for partial resolution of $^{36}\text{Ar} + \text{H}^{35}\text{Cl}$ from a $^{36}\text{Ar} + \text{H}^{35}\text{Cl} + ^{12}\text{C}_3$ sum peak at m/e 36, which is particularly significant for $^{40}\text{Ar}/^{39}\text{Ar}$ analyses of sanidine that only yield a few counts per second of ^{36}Ar .
- We have calibrated a reference gas and developed a simple analytical routine involving one peak hop, both of which are used to evaluate detector-bias correction factors that include mass dependent fractionation in the source and detector as well as detector efficiency.
- Single crystal age determinations of sanidine from the rhyolite of Alder Creek (ACs) that employ multi-collector mass spectrometry including the findings of Rivera et al. (2013) and both single crystal total fusion and incremental heating results presented here, favor an age of $1.1864 \pm 0.0003/0.0012$ Ma (95% confidence analytical/full external). This age should now serve as the preferred age for ACs because it is more accurate than the commonly used age of Nomade et al. (2005), which is more than 1% older.
- Multi-collector mass spectrometers, including the WiscAr Noblesse instrument, have led to nearly an order of magnitude improvement in analytical precision relative to traditional single-collector instruments. Independent studies using different multi-collector instruments demonstrate age heterogeneity in the Fish Canyon sanidine standard, the most widely used neutron fluence monitor in the $^{40}\text{Ar}/^{39}\text{Ar}$ community. Only a small percentage of the incrementally heated single crystals yield an age plateau, and most appear to be affected by Ar loss or contain excess Ar. These characteristics are obviously not ideal for a $^{40}\text{Ar}/^{39}\text{Ar}$ dating standard, and we suggest that use of FCs in $^{40}\text{Ar}/^{39}\text{Ar}$ -based refinement of the geologic timescale or calibration of other chronometers must be done with caution.
- High precision $^{40}\text{Ar}/^{39}\text{Ar}$ dating of single sanidine crystals with the 5-collector Noblesse is now capable of distinguishing antecrysts and xenocrysts from juvenile crystals at the 1% level of relative temporal resolution for Quaternary tuffs. Thus, $^{40}\text{Ar}/^{39}\text{Ar}$ dating of sanidine – a mineral whose daughter diffusion closure temperature is much lower than in zircon – can provide both temporal and thermal

constraints useful in leveraging the interpretation of U–Pb ID-TIMS zircon dates that commonly span tens to hundreds of thousands of years in many rhyolitic lavas and tuffs.

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Appendix A. Supplementary data

Supplementary data to this article can be found online at <http://dx.doi.org/10.1016/j.chemgeo.2016.03.024>.

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